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# integrable cases of the hamilton-Jacobi equations and dynamic systems reducible to canonical form* 

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Novel integrable cases of the Hamilton-Jacobi (HJ) equations are obtained. A method of reducing a class of non-autonomous dynamic systems to canonical form is given, and cases of their integrability are indicated. Comparison theorems are presented enabling the integrability of a dynamic system to be determined by observing the form of its Hamiltonian. The case of two bodies of variable mass in a resisting and gravitating medium are studied as an example.

1. The integration of canonical equations of motion is reduced to finding the complete integral of the corresponding HJ equation. The most interesting cases from the point of view of practical applications are the cases of integrability of the HJ, Liouville and Stackel equations $/ 1 /$ and their generalizations /2/. We shall establish new cases of integrability of the HJ equation of the form

$$
\begin{equation*}
p+\frac{1}{2} \sum_{i=1}^{n} g^{i} p_{i}^{2}+\sum_{i=1}^{n} h^{2} p_{i}-U=0 \quad\left(p=\frac{\partial V}{\partial t}, \quad p_{i}=\frac{\partial V}{\partial q_{i}}\right) \tag{1.1}
\end{equation*}
$$

which generalize the result obtained by Yarov-Yarovoi $/ 2 /$ and include the cases of integrability of Demin /3/, Liouville and stackel / //.

Theorem 1.2. If the Hamiltonian is given by the formula

$$
\begin{gather*}
H=\frac{1}{2} \frac{\gamma}{b} \sum_{i=1}^{n} \frac{1}{a_{i}\left(q_{i}\right)}\left(p_{i}-\sum_{i=1}^{k} \varphi_{i} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-  \tag{1.2}\\
\sum_{j=1}^{k} \sigma_{j} \Phi_{j}-\frac{\gamma}{b} \sum_{i=1}^{n} U_{i}\left(q_{i}\right)+\Phi_{0}(t) \\
b=\sum_{i=1}^{n} b_{i}\left(q_{i}\right), \quad \Phi_{j}=\frac{1}{\theta} \sum_{i=1}^{n} \Phi_{i j}\left(q_{i}\right) \quad(j=1,2, \ldots, k)  \tag{1.3}\\
\sigma_{j}=\varphi_{j}^{*}-c_{j} \gamma \quad(j=1,2, \ldots, k \leqslant n) \tag{1.4}
\end{gather*}
$$

where $a_{i}, b_{i}, U_{i}, \Phi_{0}, \Phi_{i j}$ are arbitrary continuous functions and $a_{i} \neq 0, b \neq 0$ and $\Phi_{i}$ are differentiable functions of the variables $q_{i}, \gamma_{*} \sigma_{j}, \varphi_{j}$ are continuous functions of time and $c_{j}$ are arbitrary constants, then the HJ equation has a complete integral

[^0]\[

$$
\begin{gather*}
V=\sum_{j=1}^{k} \varphi_{j} \Phi_{j}-\int\left(h \gamma+\Phi_{0}\right) d t+W  \tag{1.5}\\
W=\sum_{i=1}^{n} \int\left[2 a_{i}\left(U_{i}+h b_{i}-\sum_{j=1}^{k} c_{j} \Phi_{1 j}+\alpha_{i}\right)\right]^{1 / 2} d q_{i} \tag{1.6}
\end{gather*}
$$
\]

where $h$ and $\alpha_{i}$ are arbitrary constants and

$$
\begin{equation*}
\alpha_{1}+\alpha_{2}+\ldots+\alpha_{n}=0 \tag{1.7}
\end{equation*}
$$

Proof. We will seek a solution of the equation

$$
\frac{\gamma}{b} \sum_{i=1}^{n}\left[\frac{1}{2 a_{i}}\left(\frac{\partial V}{\partial q_{i}}-\sum_{j=1}^{k} \varphi_{j} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-U_{i}\right]-\sum_{j=1}^{k} \sigma_{j} \Phi_{j}+\Phi_{0}+\frac{\partial V}{\partial t}=0
$$

in the form (1.5) where $W$ is an unknown functions of $q_{1}, q_{2}, \ldots, q_{n}$. By virtue of (1.3), (1.4) we arrive at the equation

$$
\sum_{i=1}^{n}\left[\frac{1}{2 a_{i}}\left(\frac{\partial W}{\partial q_{i}}\right)^{2}-U_{i}+\sum_{j=1}^{k} c_{j} \Phi_{i j}-h b_{i}\right]=0
$$

which has the solution (1.6) where the constants $\alpha_{i}$ satisfy condition (1.7). It remains to confirm the condition

$$
\begin{equation*}
\operatorname{Det}\left\|\partial^{2} V / \partial q_{i} \partial \alpha_{j}\right\| \neq 0 \tag{1.8}
\end{equation*}
$$

We have

$$
\operatorname{Det}\left\|\frac{\partial^{2} V^{\prime}}{\partial q_{i} \partial x_{j}}\right\|=b \prod_{i=1}^{n} a_{i}\left(p_{i}-\sum_{j=1}^{k} \varphi_{j} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{-1} \neq 0
$$

since the functions $\varphi_{j}$ and $\partial \Phi_{j} / \partial q_{i}$ are continuous and according to the condition $a_{i} \neq 0, b \neq$ 0 , which completes the proof.

Corollaries. ${ }^{\circ}$. If $c_{j}=0(j=1,2, \ldots, l \leqslant k) \quad$ in (1.4), i.e. $\sigma_{j}=\varphi_{j}{ }_{j}$ and $\Phi_{j}(q)$ are arbitrary continuous functions, then the summation over $j$ in (1.6) is carried out from $j=$ $l+1$ to $j=k$.
$2^{\circ}$. If all $c_{j}=0$, then $\Phi_{0}=0, \Phi_{1}=\Phi, \Phi_{j}=0(j=2, \ldots, k)$ yields the integrable case of $/ 4 /$.
$3^{\circ}$. Putting in (1.2) all $\Phi_{j}=0$, we obtain the integrable case of $/ 2 /$, and assuming also that $\gamma=$ const and $\Phi_{0}=0$, we arrive at Liouville's theorem $/ 1 /$.

Theorem 1.2. Let $n(n+1)$ functions $\varphi_{i j}\left(q_{i}\right)$ and $U_{i}\left(q_{i}\right)(i, j=1,2, \ldots, n)$ be given, for which the determinant $\Delta=\left\|\varphi_{i j}\left(q_{i}\right)\right\| \not \equiv 0$, as well as the continuous differentiable functions $\Phi_{j}\left(q_{1}, q_{2}, \ldots, q_{n}\right)(j=1,2, \ldots, k)$ and continuous function of time $\gamma, \varphi_{j}, \sigma_{j}, \Phi_{0}(j=1,2, \ldots, k)$. Then, provided that the Hamiltonian is given by the formula

$$
\begin{align*}
H= & \frac{1}{2} \gamma \sum_{i=1}^{n}\left[\frac{A_{i}}{a_{i}}\left(p_{i}-\sum_{j=1}^{k} \varphi_{j} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-A_{i} b_{i}\right]-  \tag{1.9}\\
& \sum_{j=1}^{k} \sigma_{j} \Phi_{j}-\gamma \sum_{i=1}^{n} A_{i} U_{i}+\Phi_{0} \\
A_{i}= & \frac{1}{\Delta} \frac{\partial \Delta}{\partial \varphi_{i 1}}, \quad \Phi_{j}=\sum_{i=1}^{n} A_{i} \Phi_{i j}\left(q_{i}\right) \quad\binom{i=1,2, \ldots, n}{j=1,2, \ldots, k}  \tag{1.10}\\
& \sigma_{j}=\varphi_{j}-c_{j} \gamma \quad(j=1,2, \ldots, k \leqslant n) \tag{1.11}
\end{align*}
$$

and every coefficient $a_{i}, b_{i}, \Phi_{i j}$ depends only on the corresponding $q_{i}$ and $a_{i} \neq 0$, the HJ equation has a complete integral of the form (1.5), where

$$
\begin{equation*}
W=\sum_{i=1}^{n} \int\left[a_{i}\left(b_{i}+2 U_{i}+2 h \varphi_{i 1}-2 \sum_{j=1}^{k} c_{j} \Phi_{i j}+\sum_{j=2}^{n} \alpha_{j} \varphi_{i j}\right)\right]^{1 / 2} d q_{i} \tag{1.12}
\end{equation*}
$$

( $h, \alpha_{2}, \ldots, \alpha_{n}$ are arbitrary constants).
Proof. We will seek a solution of the corresponding equation in the form (1.5). Taking into account (1.10) and (1.11) and using the identity

$$
\sum_{i=1}^{n} \varphi_{i 1} A_{i}-1
$$

we obtain the equation

$$
\sum_{i=1}^{n} A_{i}\left[\frac{1}{a_{i}}\left(\frac{\partial U}{\partial q_{i}}\right)^{2}-b_{i}-2 U_{i}+2 \sum_{j=1}^{b} c_{j} \Phi_{i j}-2 h \varphi_{i i}\right]=0
$$

whose solution is (1.12). It remains to confirm condition (1.8). Here we will assume that $\alpha_{1}=h$. For the solution $V$ obtained we have

$$
\text { Det }\left\|\frac{\hat{\partial}^{2 V}}{\partial q_{i} \partial \alpha_{j}}\right\|=\Delta \cdot 2^{-n+1} \prod_{i=1}^{n} a_{i}\left(p_{i}-\sum_{j=1}^{n} \varphi_{j} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{-1} \neq 0
$$

Corollaries. $1^{\circ}$. If $c_{j}=0(j=1,2, \ldots, l \leqslant k)$ in (1.11), i.e. $\quad \sigma_{j}=\varphi_{j}(j=1,2, \ldots, l)$ and $\Phi_{j}(q)(j=1,2, \ldots, l)$ are arbitrary contimous functions, then the summation over $j$ in the first sum of (1.12) is carried out from $j=l+1$ to $j=k$.
$2^{\circ}$. If all $c_{j}=0$, then $\Phi_{0}=0, \Phi_{1}=\Phi, \Phi_{j}=0(j=2,3, \ldots, k)$ leads to the integrable case of /4/.

3${ }^{\circ}$. If $\Phi_{j}=0(j=1,2, \ldots, k)$ in (1.9), then we have the integrable case of $/ 2 /$. putting $\varphi_{1}=1, \varphi_{j}=0(j=2,3, \ldots, k), c_{j}=0(j=1,2, \ldots, k), \Phi_{0}=0, \gamma=$ const we arrive at the theorem due to Demin /3/ which generalizes the stảckel and Moiseyev theorems $/ 5 /$. The integrable case of stäckel / / / corresponds to $\Phi_{j} \equiv 0(j=0,1, \ldots, k), b_{i} \equiv 0, \gamma=$ const $_{*}$ a $a_{i}=$ const in formula (1.9).
2. Let the motion of a system with $n$ degrees of freedom be described by the equations

$$
\begin{align*}
q_{i}^{*} & =\partial H / \partial p_{i}, \quad p_{i}^{*}=-\partial H / \partial q_{i}+v p_{i}  \tag{2.1}\\
H & =\frac{1}{2} \sum_{i, j}^{n} g^{i j} p_{i} p_{i}+\sum_{i=1}^{n} h^{i} p_{i}-U \tag{2.2}
\end{align*}
$$

where $v(t)$ is a given function of time and $g^{i j}, h^{i}, U$ are functions of the coordinates and time, and det $\left\|g^{i j}\right\| \neq 0$. Here and henceforth the index $i$ will take the values $1,2, \ldots, n$ unless indicated otherwise.

The differential equations of a number of problems of mechanics can be reduced to the form (2.1). These inolude, in particular, systems with dissipative forces, some problems of the mechanics of controlled motion, the mechanics of bodies with variable mass and composition in the presence of reactive forces, etc. $/ 6-8 /$.

Lemma 2.1. Using the substitution

$$
\begin{equation*}
p_{i}=p_{i}^{*} \psi(t), \quad \psi(t)=\exp \left(\int v d t\right) \tag{2.3}
\end{equation*}
$$

we reduce system $(2,1)$ with Hamiltonian $(2,2)$ to the form

$$
\begin{align*}
& q_{i}^{*}=\partial H^{*} / \partial p_{i}^{*}, \quad p_{i}^{*}=-\partial H^{*} / \partial q_{i}  \tag{2.4}\\
& H^{*}\left(t, q, p^{*}\right)=\psi^{-1} H\left(t, q, p\left(p^{*}\right)\right) \tag{2.5}
\end{align*}
$$

Proof. Substituting (2.3) into (2.1) we obtain

$$
q_{i}^{*}=\partial H / \partial p_{i}, \quad p_{i}^{*}=-\psi^{\mathrm{k}} \partial H / \partial q_{i}
$$

Using (2.5) we obtain from (2.2)

$$
\begin{equation*}
\Psi^{1} \frac{\partial H}{\partial q_{i}}=\Psi \sum_{i, j=1}^{n} A_{i j} p_{i}^{*} p_{j}^{*}+\sum_{i=1}^{n} B_{i} p_{i}^{*}+\frac{c}{\psi}=\frac{\partial H *}{\partial q_{i}} \tag{2.7}
\end{equation*}
$$

where $A_{i j}(i, q), B_{i}(t, q), C(i, q)$ are the corresponding derivatives of the coefficients of the Hamiltonian (2.2) with respect to the coordinates $q_{i}$.

Let us consider the first group of equations in (2.6). Using (2.5) we obtain

$$
\begin{equation*}
\frac{\partial H^{*}}{\partial f_{i}^{*}}=\frac{\psi}{2} \sum_{i, j=1}^{n} g^{i j} p_{j}^{*}+h^{i}=\frac{\partial H}{\partial p_{i}} \tag{2.8}
\end{equation*}
$$

which proves the lemma.
According to Lemma 2.1 a system of the form (2.1) can be reduced to canonical form by substituting the moments (2.3). As a result, the Hamiltonian $H^{*}$ has the following structure:

$$
\begin{gather*}
H^{*}=\frac{1}{2} \sum_{i, j=1}^{n} g^{* i j} p_{i}^{*} p_{j}^{*}+\sum_{i=1}^{n} h^{* i} p_{i}^{*}-U^{*}  \tag{3.9}\\
g^{* i j}=\psi g^{i j}, \quad h^{* i}=-h^{i}, \quad U^{*}=\psi^{-1} U \tag{2.10}
\end{gather*}
$$

and $\operatorname{det}\left\|g^{* i j}\right\| \neq 0$.
Let the canonical system (CS) (2.4) be integrable. Then by virtue of (2.3), (2.10) we can determine in which cases the dynamic systems of the form (2.1) are integrable.

Theorem 2.1. If the CS (2.4) with the Hamiltonian $H^{*}(2.9)$ is integrable, then so is the system (2.1) with the Hamiltonian $H$ whose coefficients $g^{i j}, h^{i}, U$ are given by the formulas (2.10), and the general solution of system (2.1) has the form

$$
\begin{equation*}
\partial V / \partial \alpha_{i}=\beta_{i}, \quad p_{i}=\psi \partial V / \partial q_{i} \tag{2.11}
\end{equation*}
$$

where $V\left(t, q_{1}, q_{2}, \ldots, q_{n}, \alpha_{1}, \ldots, \alpha_{n}\right)$ is the complete integral of the HJ equation of the reduced CS (2.4) and $\alpha_{i}, \beta_{i}$ are arbitrary constants.

Thoerem 2.2. Let the CS (2.4) be autonomous and integrable. Then the non-autonomous system (2.1) will also be integrable, provided that the coefficients $g^{i j}, h^{i}, U$ and the Hamiltonian $H$ have the form

$$
\begin{equation*}
g^{i j}=\psi^{-1}(t) g^{* i j}(q), \quad h^{i}=h^{* i}(q), \quad U=\psi(t) U^{*}(q) \tag{2.12}
\end{equation*}
$$

and the general solution of the system (2.1) has the form

$$
\begin{equation*}
\partial V / \partial \alpha_{i}=\beta_{i}, \quad p_{i}=\psi \partial V / \partial q_{i}, \quad V=-\alpha_{1} t+W(q, \alpha) \tag{2.13}
\end{equation*}
$$

where $V(t, q, \alpha)$ is the complete integral of the $H J$ equation of the autonomous CS (2.4).
The proofs of the theorems 2.1 and 2.2 follow from the fact that systems of the form (2.1) can be reduced, in accordance with Lemma 2.1, to canonical form.

Let the structure of the Hamiltonian $H$ be such that $h^{i}=0$. Then the following theorem holds for systems (2.1).

Theorem 2.3. Let the CS

$$
\begin{gather*}
q_{i}^{\cdot}=\partial H / \partial p_{i}, \quad p_{i}^{*}=-\partial H / \partial q_{i}  \tag{2.14}\\
H=\frac{1}{2} \sum_{i, j=1}^{n} g^{i j} p_{i} p_{j}-U \tag{2.15}
\end{gather*}
$$

be integrable. Then the system

$$
\begin{gather*}
q_{i}^{*}=\partial H_{1} / \partial p_{i}, \quad p_{i}^{*}=-\partial H_{1} / \partial q_{i}+v p_{i}  \tag{2.16}\\
H_{1}=\frac{1}{2} \sum_{i, j=1}^{n} g^{i j} p_{i} p_{j}-\gamma(t) U \tag{2.17}
\end{gather*}
$$

will also be integrable provided that the following re\&ation holds:

$$
\begin{equation*}
v=\gamma^{*} /(2 \gamma) \tag{2.18}
\end{equation*}
$$

Proof. Let the HJ equation of the system (2.14)

$$
\begin{equation*}
\frac{1}{2} \sum_{i, j=1}^{n} g^{i j} p_{i} p_{j}-U+p=0 \quad\left(p=\frac{\partial S}{\partial t}, p_{i}=\frac{\partial S}{\partial q_{i}}\right) \tag{2.19}
\end{equation*}
$$

be integrable by the method of separation of variables and have the complete integral $S=S_{0}$ $(t, \alpha)+W(q, \alpha)$. Accoraing to Lemma 2.1 system (2.16) can be reduced to canonical form and the corresponding HJ equation has the form

$$
\begin{equation*}
\frac{1}{2} \sum_{i, j=1}^{n} g^{i j} p_{i}^{*} p_{j}^{*}-\frac{\gamma}{\psi^{2}} U+\frac{1}{\psi} p^{*}=0 \quad\left(p^{*}=\frac{\partial V}{\partial l}, p_{i}^{*}=\frac{\partial V}{\partial q_{i}}\right) \tag{2.20}
\end{equation*}
$$

If condition (2.18) holds then, by virtue of the integrability of Eq. (2.19), Eq. (2.20) whose complete integral is

$$
\begin{equation*}
V=\int \sqrt{\gamma} \frac{\partial S_{0}}{\partial t} d t+W(q, \alpha) \tag{2.21}
\end{equation*}
$$

is also integrable, and the general solution of system (2.16) has the form

$$
\begin{equation*}
\partial V / \partial \alpha_{i}=\beta_{i}, \quad p_{i}=\sqrt{\gamma} \partial W / \partial q_{i} \tag{2.22}
\end{equation*}
$$

Coroilary. If system (2.14) is autonomous: $g^{i j}=g^{i j}(q), U=U(q), H=h=$ const and integrable, and the complete integral $S=-h t+W(q, h, \alpha)$, then the non-autonomous system (2.16) is also integrable and the complete integral

$$
V=-h \int \sqrt{\gamma} d t+W(q, h, \alpha)
$$

The converse theorem can also be proved.
Theorem 2.4. Let system (2.16) with the Hamiltonian (2.15) be integrable. The the cs (2.14) with the Hamiltonian (2.17) where the quantity $\gamma$ is replaced by $1 / \gamma$ is also integrable, provided that the following relation holds:

$$
\begin{equation*}
\gamma=\exp \left(2 \int v d t\right) \tag{2.23}
\end{equation*}
$$

Proof. According to Lemma 2.1 system (2.16) can be reduced to canonical form (2.4). Let the corresponding HIT equation

$$
\begin{equation*}
\frac{\psi}{2} \sum_{i_{1}=1}^{n} g^{i j} p_{i}^{*} p_{j}^{*}-\psi^{-1} U+p^{*}=0 \quad\left(p^{*}=\frac{\partial V}{\partial t}, \quad p_{i}^{*}=\frac{\partial V}{\partial q_{i}}\right) \tag{2.24}
\end{equation*}
$$

be integrable using tho mothod of separation of variables, and possess the complete integral $V=V_{0}(t, \alpha)+W(q, \alpha)$. The HJ equation for CS (2.14) has the form

$$
\begin{equation*}
\frac{1}{2} \sum_{i, j=1}^{n} g^{i j} p_{i} p_{j}-\frac{1}{\gamma} U+p=0 \quad\left(p=\frac{\partial S}{\partial t}, \quad p_{i}=\frac{\partial S}{\partial q_{i}}\right) \tag{2.25}
\end{equation*}
$$

Let condition (2.23) hold. Then, since Eq. (2.24) is integrable, so is Eq. (2.25) whose complete integral is

$$
S=\int \exp \left(-\int v d t\right) \frac{\partial \Gamma_{0}}{\partial t} d t+W(q, \alpha)
$$

The general solution of system (2.16) has the form

$$
\partial V / \partial \alpha_{i}=\beta_{i}, \quad p_{i}=\exp \left(\int v d i\right) \partial W / \partial q_{i}
$$

and the general solution of $C S(2.14)$ is

$$
\partial S / \partial \alpha_{i}=\beta_{i}, \quad p_{i}=\partial W / \partial q_{i}
$$

Corollary. If system (2.16) can be reduced to the autonomous form $\quad g^{*^{i j}}=g^{* i j}(q)$, $U^{*}=$ $U^{*}(q), H^{*}=h=$ const, is integrable and the complete integral $V=-h t+W(q, \alpha)$, then system (2.14) is integrable and the complete integral

$$
S=-h \int \exp \left(-\int v d t\right) d t+W(q, \alpha)
$$

Theorems 2.3 and 2.4 enable us to compare the canonical systems with systems reducible to canonical form and determine their integrability from the form of the Hamiltonian. The theorems on integrability of the systems which can be reduced to canonical form given above, and the comparison theorems, together provide the basis for separating out the classes of autonomous and non-autonomous integrable systems determined by the coefficients $g^{i j}, h^{i}$, $U$ of the Hamiltonian functions and the relations (2.10) connecting them.
3. Let us consider non-autonomous dynamic systems of the form (2.1) of the Liouville and Stäckel type. Let the Hamiltonian of system (2.1) have the form

$$
\begin{equation*}
H=\frac{1}{2 b} \sum_{i=1}^{n} \frac{1}{a_{i}}\left(p_{i}-\sum_{j=1}^{k} \varphi_{j} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-\sum_{j=1}^{k} \sigma_{j} \Phi_{j}-\frac{\gamma}{b} \sum_{i=1}^{n} U_{i}+\Phi_{0} \tag{3.1}
\end{equation*}
$$

where the notation of (1.2) is used. Then the Hamiltonian of the reduced CS (2.4) will have the form

$$
\begin{equation*}
H^{*}=\frac{\psi}{2 b} \sum_{i=1}^{n} \frac{1}{a_{i}}\left(p_{i}^{*}-\sum_{j=1}^{k} \frac{\varphi_{j}}{\psi} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-\sum_{j=1}^{k} \frac{\sigma_{j}}{\psi} \Phi_{j}-\frac{\gamma}{\psi b} \sum_{i=1}^{n} U_{i}+\frac{\Phi_{0}}{\psi} \tag{3.2}
\end{equation*}
$$

Returning now to the results of Theorem 1.1 we shall point out cases when systems (2.1) with the Hamiltonian (3.1) are integrable.

Theorem 3.1. If the following conditions hold for the system (2.1) with the Hamiltonian (3.1):

$$
\begin{align*}
& b \Phi_{j}=\sum_{i=1}^{n} \Phi_{i j}\left(q_{i}\right), \quad v=\frac{\gamma^{\prime}}{2 \gamma}, \quad \sigma_{j}=\varphi_{j} \frac{d}{d t}\left(\ln \frac{\varphi_{j}}{\sqrt{\gamma}}\right)-c_{j} \gamma  \tag{3.3}\\
& (j=1,2, \ldots, k)
\end{align*}
$$

then system (2.1) is integrable and its general solution has the form

$$
\begin{equation*}
\partial V / \partial \alpha_{i}^{\prime}=\beta_{i}, \quad p_{i}=\sqrt{\gamma} \partial V / \partial q_{i} \tag{3.4}
\end{equation*}
$$

where $\alpha_{i}{ }^{\prime}, \boldsymbol{\beta}_{i}$ denote $2 n$ arbitraxy constants and

$$
\begin{equation*}
V=-\int\left(h \sqrt{\gamma}+\frac{\Phi_{0}}{\sqrt{\gamma}}\right) d t+\sum_{i=1}^{n} \int\left[2 a_{i}\left(U_{i}+h b_{i}-\sum_{j=1}^{k} c_{j} \Phi_{i j}+\alpha_{i}\right)\right]^{1 / 2} d q_{i}+\sum_{j=1}^{k} \frac{\varphi_{j}}{\sqrt{\gamma}} \Phi_{j} \tag{3.5}
\end{equation*}
$$

is the complete integral of the reduced $C S$ with the Hamiltonian (3.2), and

$$
\begin{equation*}
\alpha_{1}^{\prime}=h, \quad \alpha_{i}^{\prime}=\alpha_{i}(i=2,3, \ldots, n), \alpha_{1}=-\left(\alpha_{2}+\alpha_{3}+\ldots+\alpha_{n}\right) \tag{3.6}
\end{equation*}
$$

Theorem 3.2. Let the following conditions hold for system (2.1) and the Hamiltonian (3.1):

$$
\begin{equation*}
v=\frac{\gamma^{\cdot}}{2 \gamma}, \quad \sigma_{j}=\varphi_{j} \frac{d}{d t}\left(\ln \frac{\varphi_{j}}{\sqrt{\gamma}}\right) \quad(j=1,2, \ldots, k) \tag{3.7}
\end{equation*}
$$

Then system (2.1) will be integrable, its general solution will have the form (3.4) and the complete integral of the reduced $C S$ with the Hamiltonian (3.2) will have the form (3.5) where all $c_{j} \equiv 0(j=1,2, \ldots, k)$ and relations (3.6) al so hold.

The proof of the theorems follows from Lemma 2.1 and Theorem 1.1.
Let the Hamiltonian of system (2.1) have the form

$$
\begin{equation*}
H=\frac{1}{2} \sum_{i=1}^{n}\left[\frac{A_{i}}{a_{i}}\left(p_{i}-\sum_{j=1}^{k} \varphi_{j} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-\gamma A_{i} b_{i}\right]-\sum_{j=1}^{k} \sigma_{j} \Phi_{j}-\gamma \sum_{i=1}^{n} A_{i} U_{i}+\Phi_{0} \tag{3.8}
\end{equation*}
$$

where the notation of ( 1.9 ) is used. Then the Hamiltonian of the reduced $C S$ will be equal to

$$
\begin{equation*}
H^{*}=\frac{\psi}{2} \sum_{i=1}^{n}\left[\frac{A_{i}}{a_{i}}\left(p_{i}^{*}-\sum_{j=1}^{k} \frac{\varphi_{j}}{\psi} \frac{\partial \Phi_{j}}{\partial q_{i}}\right)^{2}-\frac{\gamma}{\psi^{2}} A_{i} b_{i}\right]-\sum_{j=1}^{k} \frac{\sigma_{j}}{\psi} \Phi_{j}-\frac{\gamma}{\psi} \sum_{i=1}^{n} A_{i} U_{i}+\frac{\Phi_{0}}{\psi} \tag{3.9}
\end{equation*}
$$

The following theorems hold for systems (2.1) with the Hamiltonian (3.8).
Theorem 3.3. Let system (2.1) have a Hamiltonian $H$ of the form (3.8). Then, provided that the conditions

$$
\begin{align*}
& \Phi_{j}=\sum_{i=1}^{n} A_{i} \Phi_{i j}\left(q_{i}\right), \quad v=\frac{\gamma^{\circ}}{2 \gamma}, \quad \sigma_{j}=\varphi_{j} \frac{d}{d t}\left(\ln \frac{\varphi_{j}}{\sqrt{\gamma}}\right)-c_{j} \gamma  \tag{3.10}\\
& (j=1,2, \ldots, k)
\end{align*}
$$

hold, system (2.1) will be integrable and its general solution will have the form

$$
\begin{equation*}
\partial V / \partial \alpha_{i}=\beta_{i}, \quad p_{i}=\sqrt{\gamma} \partial V / \partial q_{i}, \quad \alpha_{1}=h \tag{3.11}
\end{equation*}
$$

$$
\begin{align*}
V= & -\int\left(h \sqrt{\gamma}+\frac{\Phi_{0}}{\sqrt{\gamma}}\right) d t+\sum_{i=1}^{n} \int\left[a _ { i } \left(b_{i}+2 U_{i}+2 h \varphi_{i 1}-\right.\right.  \tag{3.12}\\
& \left.\left.2 \sum_{j=1}^{h} c_{j} \Phi_{i j}+\sum_{j=2}^{n} \omega_{j} \varphi_{i j}\right)\right]^{1 / t} d q_{i}+\sum_{j=1}^{h} \frac{\varphi_{j}}{\sqrt{\gamma}} \Phi_{j}
\end{align*}
$$

is the complete integral of the reduced CS with the Hamiltonian (3.9).
Theorem 3.4. Let system (2.1) have a Hamiltonian $H$ of the form (3.8). Then, provided that conditions (3.7) hold, system (2.1) will be integrable, its general solution will have the form (3.11) and the completc integral of the reduced CS with the Hamiltonian (3.9) will have the form (3.12) where all $c_{j} \equiv 0(j=1,2, \ldots, k)$.

The proofs of the theorems follow from Lemma 2.1 and Theorem 1.2.
4. Examples. $1^{\circ}$. We shall illustrate the general method of reduction to canonical form and integration of dynamic systems by considering the problem of two bodies (material points) of variable mass, which may find application in celestial mechanics $/ 9 /$. The bodies attract each other in accordance with Newton's Law and are situated within a gaseous or dust cloud exerting additional "frictional" forces and Hooke's elastic force. The equations of motion have the form

$$
\begin{equation*}
\mathbf{r} \cdot \ddot{=}=-\mu(t) r^{-3} \mathbf{r}+v(t) \mathbf{r}^{*}+x(t) \mathbf{r}, \quad \mu(t)=G M(t) \tag{4.1}
\end{equation*}
$$

where $G$ is the gravitational constant, $M(t)$ is the mass of the two bodies, $v, x$ are continuous functions of time characterizing the background, and $r$ is the radius vector of the motion of one material point relative to the other. Using the spherical coordinates $r, p, i$ we can write the equations of motion (4.1) in the form (2.1), using the results of Lemma 2.1 . The corresponding $H J$ equation will have the form

$$
\begin{equation*}
\frac{\psi}{2}\left[\left(\frac{\partial V}{\partial r}\right)^{2}+\frac{1}{r^{2}}\left(\frac{\partial r}{\partial \varphi}\right)^{2}+\frac{1}{r^{2} \cos ^{2} \varphi}\left(\frac{\partial V}{\partial \lambda}\right)^{2}\right]-\frac{\mu}{\psi r}-\frac{\kappa}{2 \varphi} r^{2}+\frac{\partial V}{\partial t}=0 \tag{4.2}
\end{equation*}
$$

Let the conditions

$$
\begin{equation*}
\mu / \psi^{2}=\mu_{0}, \quad \kappa / \psi^{2}=\chi_{0}, \quad v=\mu^{\cdot} /(2 \mu) \tag{4.3}
\end{equation*}
$$

hold. Then Eq. (4.2) can be integrated and its complete integral will have the form

$$
\begin{align*}
& \boldsymbol{V}=-\alpha_{1} \int\left(\frac{\mu}{\mu_{0}}\right)^{1 / 3} d t+\int\left[2\left(\frac{\mu_{0}}{r}+\frac{\alpha_{0}}{2} r^{2}\right)-\frac{\alpha_{2}^{2}}{r^{2}}+2 \alpha_{1}\right]^{1 / 2} d r+  \tag{4.4}\\
& \int\left[\alpha_{2}^{2}-\frac{\alpha_{3}^{2}}{\cos ^{2} \varphi}\right]^{1 / 2} d \varphi+\alpha_{3} \lambda
\end{align*}
$$

where $\alpha_{1}, \alpha_{2}, \alpha_{3}$ are arbitrary constants. The general solution of problem (4. 1 ) will be given by formulas of the form (2.11).

We note that problem (4.1) is integrable when the potential is of a more general type

$$
\begin{equation*}
U(r, \varphi, \lambda, t)=\mu(t)\left[f(r)+\frac{\Phi(\Phi)}{r^{2}}+\frac{\Phi(\lambda)}{r^{2} \cos ^{2} \varphi}\right] \tag{4.5}
\end{equation*}
$$

where $f(r), \Phi(\varphi), \Phi(\ell)$ are arbitrary differential functions. Carrying out the same arguments for problem (4.1) with the potential (4.5), we obtain the complete integral $v$ of the corresponding $H J$ equation

$$
\begin{align*}
& V=-\alpha_{1} \int\left(\frac{\mu}{\mu_{9}}\right)^{1 / 2} d t+\int\left[2 \mu_{0} f(r)-\frac{\alpha_{2}^{2}}{r^{2}}+2 x_{1}\right]^{1 / 2} d r+  \tag{4,6}\\
& \int\left[\alpha_{2}^{2}-2 \mu_{0} \Phi(\varphi)-\frac{\alpha_{3}^{2}}{\cos ^{2} \varphi}\right]^{1 / 2} d \varphi+\int\left[\alpha_{3}^{2}-2 \mu_{0} \Phi(\lambda)^{1 / 2} d \lambda\right.
\end{align*}
$$

The general solution of the problem will be given by formulas (2.11) whexe the complete integral $V$ has the form (4.6). The solution obtained generalizes the results of /9/ to the case of a potential of gencral type (4.5) for a gravitating and resisting medium.
$2^{\circ}$. Let us consider the motion of a passively gravitating point in a gravitational field of two bodies $M_{1}(t)$ and $M_{2}(t)$ of variable mass, moving along a straight line passing through the centres of mass. We take into account the influcnce of the gravitating medium generating additional forces acting on the material point, analogous to "frictional" forces, and Hooke's elastic force. The motion of the attracting bodies themselves is determined by the Gil'denMeshcherskii problem which has linear solutions /10/. Let the masses $M_{i}$ and $M_{2}$ vary with time
according to the same law.
The equations of motion of a point in a rectangular system of coordinates with origin at the centre of mass of $M_{1}$ and $M_{2}$ and $x$ axis collinear with the line of motion of bodies with finite masses, will be written in the form

$$
\begin{equation*}
\mathbf{r}^{\bullet}=\operatorname{grad} U+\boldsymbol{v} \mathbf{r}^{\cdot}+\beta \mathbf{r} \tag{4,7}
\end{equation*}
$$

where $\mathbf{r}$ is the radius vector of the point, $\nu(t), \beta(t)$ are continuous functions determining the parameters of the gravitating medium, the potential has the form

$$
U=\mu_{1} / r_{1}+\mu_{2} / r_{2}, \quad \mu_{i}=G M_{i}(t) \quad(i=1,2)
$$

and $r_{1}, r_{2}$ are the distances between the material point and the attracting bodies.
Using the notation of $/ 4 /$, we shall write Eqs. (4.7) in the form (2.16) where $i=1,2,3$, and the Hamiltonian has the form

$$
\begin{aligned}
H= & \frac{2}{r_{12}^{2}\left(\lambda^{2}-\eta^{2}\right)}\left\{\left(\lambda^{2}-1\right)\left[p_{\lambda}-\frac{r_{12}^{*} r_{12}}{4} \frac{\partial v}{\partial \lambda}\right]^{2}+\right. \\
& \left.\left(1-\eta^{2}\right)\left[p_{\eta}-\frac{r_{12} \cdot r_{12}}{4} \frac{\partial v}{\partial \bar{\eta}}\right]^{2}+\left(\frac{1}{\lambda^{2}-1}+\frac{1}{1-\eta^{2}}\right) p_{w}^{2}\right\}- \\
& \frac{r_{12} \cdot 2}{4} v-2 \frac{\left(\mu_{1}+\mu_{2}\right) \lambda-\left(\mu_{1}-\mu_{2}\right) \eta}{r_{12}\left(\lambda^{2}-\eta^{2}\right)}-\frac{\beta r_{12}^{2}}{4} v \\
v= & \frac{\lambda^{2}+\eta^{2}}{2}+\lambda \eta k+\frac{k^{2}-1}{2}
\end{aligned}
$$

and the coordinates $\lambda, \eta, w$ are regarded as generalized coordinates $q_{i}(i=1,2,3)$. According to Lemma 2.1 system (2.16) can be reduced to the canonical form

$$
\begin{align*}
& q_{i}^{*}=\partial H^{*} / \partial p_{i}^{*}, \quad p_{i}^{*}=-\partial H^{*} / \partial q_{i} \quad(i=1,2,3)  \tag{4.8}\\
& H^{*}\left(t, q, p^{*}\right)=中^{-1} H\left(t, q, p\left(p^{*}\right)\right)
\end{align*}
$$

Let the following conditions hold:

$$
\begin{align*}
& q=\frac{r_{12} \cdot r_{12}}{4 \psi}, \quad \frac{r_{12} \mu}{\psi^{2}}=C_{6}, \quad \frac{r_{122} \mu_{i}}{\psi^{2}}=C_{i} \quad(i=1,2)  \tag{4.9}\\
& r_{12}^{3}\left(r_{12} \cdot \bullet-v r_{12} \cdot-\beta r_{12}\right) /\left(8 \psi^{2}\right)=C ; \quad C=0, \quad k \neq 0 ; \quad C \neq 0, \quad k=0
\end{align*}
$$

The Gil'den-Meshcherskii problem determining the motion of the bodies $M_{1}$ and $M_{a}$ admits of the following linear solutions /lo/:

$$
\begin{align*}
& \mathbf{r}_{12}=\mu^{-1} \lambda_{\lambda_{0}}, \quad \lambda_{0}^{3}=-\mu_{0} / b_{0}, \quad b_{0} \neq 0, \quad b_{1}=0  \tag{4.10}\\
& \mathbf{r}_{12}=\left(3 b_{1} \zeta\right)^{2 / 3 \mu-1 \lambda_{0}}, \quad b_{1} \neq 0 ; \quad \zeta=\int \mu^{2} d t
\end{align*}
$$

Assuming $\mu(t)$ to be given, we can determine the function $v(t), \beta(t), \varphi(t), \psi(t)$, according to /lo/, from the conditions of integrability of (4.9), taking (4.10) into account. As a result we obtain for $b_{1}=0$

$$
\begin{align*}
& r_{12}=\mu^{-1} \lambda_{0}, \quad v=0, \quad \psi=\sqrt{\lambda_{0} / C_{0}}  \tag{4.11}\\
& \beta=\left(b_{0}-8 C C_{0}^{-1} \lambda_{0}^{-3}\right) \mu^{4}, \quad \varphi=-{ }^{1 / 4} \sqrt{C_{0}} \lambda_{0}^{3 /}{ }^{4} \cdot \mu^{-3}
\end{align*}
$$

and for $b_{1} \neq 0$

$$
\begin{align*}
& r_{12}=\left(3 b_{1} \zeta\right)^{2 / 3} \mu^{-1} \lambda_{0}, \quad v=1_{3} \mu_{5}^{2}-1, \quad \psi=\sqrt{\lambda_{0}^{\prime} / C_{0}}\left(3 b_{15}^{5}\right)^{1 / 3}  \tag{4.12}\\
& \beta=\left(-2 b_{1}^{2}+b_{0}-8 C C_{0}^{-1} \lambda_{0}^{-3}\right)\left(9 b_{1}^{2}\right)^{-1} \mu^{4} \zeta^{-2}+v \mu^{\cdot} \mu^{-1} \\
& \varphi=1 / 4 \sqrt{C_{0}}\left[2 b_{1}-\mu^{\prime} \mu^{-3} 3 b_{15}^{5}\right] \lambda_{0}^{3 / 2}
\end{align*}
$$

Formulas (4.11) and (4.12) determine two classes of solutions of the bounded linear problem of three bodies of variable mass, taking into account the gravitating and resistive medium.

The complete integral of the HJ equation of the $C S$ (4.8) has the form

$$
\begin{aligned}
& V=\varphi v-h \int \frac{\psi}{r_{12}^{2}} d t+\int \frac{\sqrt{Q_{+}(\lambda)}}{\lambda^{2}-1} d \lambda+\int \frac{\sqrt{Q_{-}(\eta)}}{1-\eta^{2}} d \eta+\alpha_{3} w \\
& Q_{ \pm}(\lambda)=\left(\lambda^{2}-1\right)\left[-\frac{C}{2} \lambda^{4}+\frac{h+C}{2} \lambda^{2}+\left(C_{1} \pm C_{2}\right) \lambda+\alpha_{2}\right]-\alpha_{3}^{2}
\end{aligned}
$$

and $C=0$ when $k \neq 0 ; C \neq 0, C_{1}=C_{2}$ when $k=0 ; h, \alpha_{3}, \alpha_{3}$ are arbitrary constants. The functions $\varphi, r_{12}, \psi$ are given by the formulas (4.11) or (4.12).

The general solution of the problem has the form ( $\beta_{i}$ are arbitrary constants)

$$
\partial V / \partial \alpha_{i}-\beta_{i}, \quad p_{i}-\psi(t) \partial V \partial q_{i} \quad(i=1,2,3), \quad \alpha_{1}=h
$$

Problem (4.7) was studied for the case when $v=0, \beta=0$ in /4, 11/. A spatial-temporal transformation method was used in /11/ to show that the problem can be solved when the mass $\mu(t) \quad$ varies according to the first Meshcherskii law, and in the case of the generalized Meshcherskii law it has a solution when the masses are equal to each other $\mu_{1}=\mu_{2}$. A solution of the problem was obtained in /4/using the Jacobi method for the case when the mass change obeys the first Meshcherskii law.

We shall now state the following, more general result. Assuming that the functions $\mu(t), r_{12}(t)$ are given, we obtain from the conditions of integrability (4.9) the expressions

$$
\begin{align*}
& v(t)=1 / 2\left(r_{12}^{*} / r_{12}+\mu^{\cdot} / \mu\right)  \tag{4.13}\\
& \beta(t)=r_{12}^{*} / r_{12}-1 / 2\left(r_{12}^{*} / r_{12}+\mu^{\cdot} / \mu\right) r_{12}^{*} / r_{12}-8 C C_{0}^{-1} \mu / r_{12}^{3}
\end{align*}
$$

and in (4.13) $C=0$ when $k \neq 0$. This implies that the equations of motion (4.7) can be integrated for any specified, continuously differentiable functions $\quad \mu(t), r_{12}(t)$ which do not vanish within the time interval in question, provided that $v(t), \beta(t)$ can be found from formulas (4.13) .

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